

### Sulfur reactions and photodissociations for Titan:

This chemical scheme for sulfur species has been used in the following paper:

K.M. Hickson, J.C. Loison, T. Cavalié, E. Hébrard, and M. Dobrijevic. The evolution of infalling sulfur species in Titan's atmosphere. *Astronomy and Astrophysics*. 2014.

See KIDA for the nomenclature of rate constants and uncertainties (See <http://kida.obs.u-bordeaux1.fr/help>).

Reaction	$\Delta H_r(298K)$ kJ/mol	$\alpha$	$\beta$	$\gamma$	$F_0$	$g$	references
H + SH → H <sub>2</sub> + S	-79	3.0e-11	0	0	5	0	(Cupitt & Glass 1975, Peng <i>et al.</i> 1999)
H + H <sub>2</sub> S → H <sub>2</sub> + SH		3.66e-12	1.94	455	1.6	100	(Peng <i>et al.</i> 1999)
H + CS → HCS		k <sub>0</sub> =7.5e-34 k <sub>∞</sub> =1.0e-10	0.2 0	0 800	3 3	0 100	/ H + CO, barrier calculated at M06-2X/cc-pVTZ
H + HCS → H <sub>2</sub> + CS		1.0e-10	0	0	2	0	/ H + HCO and also (Galland <i>et al.</i> 2001)
H + H <sub>2</sub> CS → HCS + H <sub>2</sub>		3.3e-11	0	1400	3	200	/ H + H <sub>2</sub> CO with smaller barrier
H + CH <sub>3</sub> S → H <sub>2</sub> CS + H <sub>2</sub> → CH <sub>3</sub> + SH	-210 -50	1.0e-11 3.0e-11	0 0	0 0	3 3	0 0	by comparison with CH <sub>3</sub> O + H considering than the H <sub>2</sub> CS + H <sub>2</sub> channel is much less exothermic than the corresponding H <sub>2</sub> CO + H <sub>2</sub> (-210 compared to -344).
H + CH <sub>3</sub> SH → CH <sub>3</sub> S + H <sub>2</sub>		7.8e-11	0	1310	1.8	200	(Amano <i>et al.</i> 1983)
<sup>2</sup> H + <sup>2</sup> HCCS → H <sub>2</sub> CCS → H <sub>2</sub> + CCS → <sup>3</sup> CH <sub>2</sub> + CS	-373 -1.0 +102	k <sub>0</sub> =2.0e-27 k <sub>∞</sub> =1.0e-10 0 0	-1.5 0	0 0	3 3	0 0	M06-2X/cc-pVTZ / Semi empirical model
<sup>2</sup> H + <sup>2</sup> CCS → HCCS → C <sub>2</sub> H + S → CH + CS	-425 +52 +106	k <sub>0</sub> =1.0e-27 k <sub>∞</sub> =1.0e-10 0 0	-1.5 0	0 0	3 3	0 0	M06-2X/cc-pVTZ / Semi empirical model
H + CH <sub>3</sub> CS → HCS + CH <sub>3</sub> → H <sub>2</sub> + H <sub>2</sub> CCS		4.0e-11 2.0e-11	0 0	0 0	3 3	0 0	/ H + CH <sub>3</sub> CO
H + CH <sub>3</sub> CHS → H <sub>2</sub> + CH <sub>3</sub> CS	-17	2.23e-11	0	1200	3	200	/ H + CH <sub>3</sub> CHO with smaller E#
H + C <sub>3</sub> S → HC <sub>3</sub> S		k <sub>0</sub> =3.0e-25	-2.93	176	10	0	Equal to H + C <sub>4</sub> H <sub>2</sub>

		$k_{\infty}=1.0e-10$	0	800	3	100	Equal to H + CS
H + HC <sub>3</sub> S	→ H <sub>2</sub> + C <sub>3</sub> S → C <sub>2</sub> H <sub>2</sub> + CS	-118 -177	2.0e-11 2.0e-11	0 0	3 3	0 0	Energy calculated at M06-2X/cc-pVTZ, the main exit channel may be H <sub>2</sub> C <sub>3</sub> S formation.
H + C <sub>2</sub> H <sub>2</sub> S	→ C <sub>2</sub> H <sub>2</sub> + SH		2.87e-11	0	946	3	/ by comparison with C <sub>2</sub> H <sub>4</sub> S + H
H + C <sub>2</sub> H <sub>4</sub> S	→ C <sub>2</sub> H <sub>4</sub> + SH		2.87e-11	0	946	1.8	(Lee <i>et al.</i> 1977a)
H + OCS	→ SH + CO		9.06e-12	0	1940	2	200 (Lee <i>et al.</i> 1977b, Rice <i>et al.</i> 1993a, Tsunashima <i>et al.</i> 1975, Adriaens <i>et al.</i> 2010)
H + N <sub>2</sub> S	→ N <sub>2</sub> + SH	-270	2.1e-10	0	2600	1.6	100 M06-MP2/cc-pVTZ calculations
H + NCS	→ HCN + S → HNC + S	-93 -40	0 1.0e-10	0	0	2	0 submerged exit barrier on the triplet surface, need ISC as there is likely a barrier on the triplet entrance barrier but not on the entrance singlet surface by comparison with H + NCO (Klippenstein & Harding 2009))
C + H <sub>2</sub> S	→ H + HCS	-178	2.1e-10	0	0	1.6	0 (Deeyamulla & Husain 2006, Galland <i>et al.</i> 2001)
C + OCS	→ CO + CS	-386	5.6e-11	0	0	2	0 (Loison <i>et al.</i> 2012b, Deeyamulla & Husain 2006, Deeyamulla & Husain 2007)
C + CS <sub>2</sub>	→ CS + CS	-242	1.6e-10	0	0	3	0 (Deeyamulla & Husain 2006)
<sup>2</sup> CH + <sup>1</sup> CS	→ <sup>2</sup> H + <sup>3</sup> C <sub>2</sub> S → <sup>3</sup> S + <sup>2</sup> C <sub>2</sub> H	-106 -54	1.5e-10 5.0e-11	0 0	0 0	2 2	0 Very likely no barrier
<sup>2</sup> CH + <sup>1</sup> H <sub>2</sub> S	→ <sup>2</sup> H + <sup>1</sup> H <sub>2</sub> CS	-239	2.0e-10	0	0	2	0 By comparison with CH + H <sub>2</sub> O (Bergeat <i>et al.</i> 2009, Hickson <i>et al.</i> 2013)
<sup>2</sup> CH + <sup>1</sup> OCS	→ <sup>2</sup> HCS + <sup>1</sup> CO → <sup>2</sup> H + <sup>1</sup> CS + <sup>1</sup> CO	-268 -57	4.0e-10 0	0	0	2	0 (Loison <i>et al.</i> 2012b)
<sup>2</sup> CH + <sup>1</sup> CS <sub>2</sub>	→ <sup>2</sup> HCS + <sup>1</sup> CS	-124	4.0e-10	0	0	2	0 Considering the high CH reactivity and the stability of HCS
CH + N <sub>2</sub> S	→ N <sub>2</sub> + HCS → HCN + NS	-497 -364	1.5e-11 1.5e-11	0 0	-257 -257	2 2	0 /CH + N <sub>2</sub> O (Becker <i>et al.</i> 1993)
<sup>1</sup> CH <sub>2</sub> + CS	→ HCCS + H	-158	2.0e-10	0	0	2	0 Very likely no barrier
<sup>3</sup> CH <sub>2</sub> + CS	→ HCCS + H	-102	6.0e-11	0	0	2	0 no barrier at M06-2X/cc-pVTZ and also at MP2/cc-pVTZ level
CH <sub>2</sub> + CCS	→ HC <sub>3</sub> S + H		1.0e-10	0	0	3	0 / CH <sub>2</sub> + CH <sub>3</sub>

CH <sub>2</sub> + HCCS → H <sub>2</sub> C <sub>3</sub> S + H	-179	1.0e-10	0	0	3	0	/ CH <sub>2</sub> + CH <sub>3</sub>
<sup>3</sup> CH <sub>2</sub> + C <sub>3</sub> S → 1-C <sub>3</sub> H <sub>2</sub> + CS → HC <sub>4</sub> S + H	-108 -186	6.0e-11 0	0	0	2	0	Assuming no barrier / <sup>3</sup> CH <sub>2</sub> + CS. HC <sub>4</sub> S very likely the main product but should react with H (and also to CH <sub>3</sub> ) to give C <sub>3</sub> H <sub>2</sub> + CS. We avoid introducing new minor species.
CH <sub>2</sub> + NS → H + HCNS(=HNCS)	-193	1.0e-10	0	0	3	0	/ CH <sub>2</sub> + NO at 150K
CH <sub>3</sub> + CS → H <sub>2</sub> CCS + H	-30	2.0e-12	0	2400	3	400	Average between M06-2X and MP2 /cc-pVTZ
CH <sub>3</sub> + SH → CH <sub>3</sub> SH  → H <sub>2</sub> CS + H <sub>2</sub> → HCSH + H <sub>2</sub> → CH <sub>2</sub> SH + H → CH <sub>3</sub> S + H → <sup>1</sup> CH <sub>2</sub> + H <sub>2</sub> S → CH <sub>4</sub> + S	-308  -160 +27 +86 +50 +122 -82	k <sub>0</sub> =1.0e-28 k <sub>r</sub> =1.2e-10 0 0 0 0 0	-3.5 0	0 0	3 3	0 0	M06-2X/cc-pVTZ/semi empirical model  TS(→ H <sub>2</sub> CS + H <sub>2</sub> ) very likely > 0 TS(→ HCSH + H <sub>2</sub> ) : +81 kJ/mol
CH <sub>3</sub> + HCS → CH <sub>4</sub> + CS		9.3e-11	0	0	2	0	/ CH <sub>3</sub> + HCO. We neglect the association reaction as CH <sub>3</sub> CHS absorbs strongly in the 190-230 nm (Rosengren 1962) and dissociate leading very likely to HCS, CS and CH <sub>3</sub> CS all leading to CS formation.
CH <sub>3</sub> + C <sub>2</sub> S → H + H <sub>2</sub> C <sub>3</sub> S → C <sub>2</sub> H <sub>3</sub> + CS	-159 -126	2.0e-11 2.0e-11	0 0	0 0	3 3	0 0	the TS from <sup>2</sup> CH <sub>3</sub> CCS toward <sup>2</sup> CH <sub>2</sub> CHCS is found -218 kJ/mol below the entrance level followed by a small barrier equal to 6.7 kJ/mol above the exit channel for H elimination, and there is no exit TS at M06 and MP2 level for the <sup>2</sup> C <sub>2</sub> H <sub>3</sub> + <sup>1</sup> CS production
CH <sub>3</sub> + HCCS → C <sub>2</sub> H <sub>4</sub> + CS	-151	2.0e-11	0	0	3	0	See text
CH <sub>3</sub> + NS → H + H <sub>2</sub> CNS → H <sub>3</sub> CNS (SOOT)	+21	0 k <sub>0</sub> =1.1e-29 k <sub>r</sub> =1.1e-11	-3.5 0.60	0 0	2 1.6	0 0	/ CH <sub>3</sub> + NO
C <sub>2</sub> + H <sub>2</sub> S → C <sub>2</sub> H + HS	-226	2.0e-10	0	0	2	0	(Kaiser <i>et al.</i> 2002, Wang <i>et al.</i> 2003)
C <sub>2</sub> + CS <sub>2</sub> → C <sub>2</sub> S + CS	-198	2.0e-10	0	0	2	0	By comparison with C <sub>2</sub> (a <sup>3</sup> Π <sub>u</sub> ) (Hu <i>et al.</i> 2014, Huang <i>et al.</i> 2004) reaction, the ground C <sub>2</sub> (X <sup>1</sup> Σ <sub>g</sub> <sup>+</sup> ) being more reactive (Canosa <i>et al.</i>

								2007, Daugey <i>et al.</i> 2008).
$C_2 + OCS \rightarrow C_2S + CO$	-342	2.0e-10	0	0	2	0		By comparison with $C_2 + CS_2$
${}^2C_2H + {}^1CS \rightarrow {}^1C_3S + {}^2H$	-62	1.0e-10	0	0	2	0		/ $C_2H + CO$ (Lander <i>et al.</i> 1990, Petrie 1996)
${}^2C_2H + {}^1H_2S \rightarrow {}^1C_2H_2 + {}^2HS$	-172	1.0e-10	0	0	3	0		No barrier for preliminary DFT calculations
${}^2C_2H + {}^1CS_2 \rightarrow HCCS + CS$	-25	6.0e-11	0	0	6	0		No barrier in the entrance valley at M06-2X/cc-pVTZ level but may be a small one in the exit valley.
${}^2C_2H + {}^1OCS \rightarrow HCCS + CO$	-169	4.0e-11	0	0	5	0		Likely no barrier
${}^2C_2H + {}^1C_3S \rightarrow \text{products}$		1.0e-10	0	0	3	0		/ $C_2H + CO$ (Lander <i>et al.</i> 1990, Petrie 1996)
$C_2H_3 + HS \rightarrow C_2H_2 + H_2S$	-222	5.0e-11	0	0	2	0		$C_2H_3 + OH$ (Tsang & Hampson 1986)
$\rightarrow H_2C_2 + H_2S$	-30	0						HSCCH likely a minor product by comparison
$\rightarrow H_2CCSH + H$	+83	0						with $C_2H_3CN$ photodissociation ( $\ll 50\%$ )
$\rightarrow HSCCH + H_2$	-177	0						neglected here.
$C_2H_3 + CS \rightarrow H_2C_3S + H$	-33	3.0e-11	0	0	3	0		no barrier at M06-2X and MP2 level /cc-pVTZ
$C_2H_3 + C_3S \rightarrow H_2C_5S + H$	-97	3.0e-11	0	0	3	0		Equal to $C_2H_3 + CS$
$CN + H_2S \rightarrow HCN + HS$	-165	2.0e-10	0	0	2	0		Very likely no barrier as $H_2S$ is much more reactive than $H_2O$
$CN + CS_2 \rightarrow NCS + CS$	+8	0						
$CN + OCS \rightarrow NCS + CO$	-136	1.0e-10	0	0	3	0		(Loison <i>et al.</i> 2012b)
$O + CS \rightarrow S + CO$		2.61e-10	0	758	2	100		(Lilenfeld & Richardson 1977)
$OH + CS \rightarrow H + OCS$		1.7e-10	0	0	3	0		(Rice & Chabalowski 1994, Loison <i>et al.</i>
$\rightarrow CO + HS$		3.0e-11	0	0	3	0		2012a)
$S + CH_3 \rightarrow H_2CS + H$	-87	1.4e-10	0	0	3	0		by comparison with $O + CH_3$
$\rightarrow H_2 + CS + H$	+93	0						
$S + C_2H_2 \rightarrow C_2H_2S$		$k_s = 1.4e-11$ $k_0 = 5.5e-30$	0	1010	1.8	200		(Strausz <i>et al.</i> 1971, Little & Donovan 1973)
			0	0	10	0		= $k_0(OH + C_2H_2)$
$S + C_2H_3 \rightarrow SH + C_2H_2$	-199	1.0e-11	0	0	2	0		By comparison with $O + C_2H_3$ (Heinemann <i>et al.</i>
$\rightarrow H + H_2C_2S$	-161	6.0e-11	0	0	2	0		1988, Harding <i>et al.</i> 2005)
$S + C_2H_4 \rightarrow C_2H_4S$		$k_s = 7.14e-12$ $k_0 = 1.0e-28$	0	795	1.8	200		(Davis <i>et al.</i> 1972)
			0	0	10	0		/ $OH + C_2H_4$
$S + CO \rightarrow OCS$		0						(Loison <i>et al.</i> 2012b)

S + NH	→ NS + H	-116	6.6e-11	0	0	2	0	/ O + NH
S( <sup>1</sup> D) + N <sub>2</sub>	→ S( <sup>3</sup> P) + N <sub>2</sub>	-111	8.5e-11	0	0	1.6	0	(Black & Jusinski 1985)
S( <sup>1</sup> D) + H <sub>2</sub>	→ SH + H	-27	2.1e-10	0	0	1.6	0	(Black & Jusinski 1985)
S( <sup>1</sup> D) + CH <sub>4</sub>	→ SH + CH <sub>3</sub>	-27	1.7e-10	0	0	2.0	0	(Black & Jusinski 1985)
	→ H <sub>3</sub> CS + H	+20	0					
	→ H <sub>2</sub> CS + H <sub>2</sub>	-200	0					
SH + CO	→ HSCO		0					We neglect this reaction as there is a barrier in the entrance and the well is small (Rice <i>et al.</i> 1993b, Rice & Chabalowski 1994)
SH + CS	→ H + CS <sub>2</sub>	-103	2.0e-10	0	0	3	0	/ OH + CS
<sup>2</sup> N + <sup>1</sup> C <sub>3</sub> S	→ <sup>2</sup> CN + <sup>3</sup> C <sub>2</sub> S	-309	2.0e-11	0	0	3	0	/ <sup>2</sup> N + <sup>1</sup> CO (Herron 1999)
	→ <sup>2</sup> C <sub>2</sub> N + <sup>1</sup> CS	-355	2.0e-11	0	0	3	0	
<sup>2</sup> N + N <sub>2</sub> S	→ N <sub>2</sub> + NS		2.0e-11	0	0	3	0	<sup>2</sup> N + N <sub>2</sub> O / (Herron 1999) with no barrier
<sup>2</sup> N + <sup>1</sup> CS	→ <sup>2</sup> CN + <sup>3</sup> S	-257	4.0e-11	0	0	3	0	/ <sup>2</sup> N + <sup>1</sup> CO (Herron 1999)
<sup>4</sup> N + <sup>2</sup> HCCS	→ <sup>1</sup> HCN + <sup>3</sup> CS	-65	8.0e-11	0	0	3	0	/ N + CH <sub>3</sub>
<sup>2,4</sup> N + <sup>1</sup> SH	→ <sup>2</sup> NS + <sup>2</sup> H	-94	4.5e-11	0	0	2	0	/ N + OH but NS is easily photodissociated
<sup>2,4</sup> N + NS	→ N <sub>2</sub> + S		4.0e-11	0	0	2	0	/ N + NO
<sup>2,4</sup> N + HCS	→ NH + CS	-114	2.0e-11	0	0	3	0	
	→ H + NCS	-261	1.0e-11	0	0	3	0	
	→ HCN + S	-354	4.0e-11	0	0	3	0	
NH <sub>2</sub> + CS	→ H + HNCS	-147	6.0e-11	0	0	3	0	No barrier at M06-2X/cc-pVTZ level and very likely also at MP2/cc-pVTZ level
NH + NS	→ N <sub>2</sub> S + H		2.9e-11	-0.30	-77	1.6	0	/ NH + NO
	→ N <sub>2</sub> + SH		1.2e-11	-0.30	-77	1.6	0	/ NH + NO
NH + CS	→ H + NCS	-147	2.0e-11	0	0	3	0	No barrier at M06-2X/cc-pVTZ level and at
	→ HNC + S	-187	4.0e-11	0	0	3	0	MP2/cc-pVTZ level
NS + NS	→ N <sub>2</sub> S + S	-123	4.0e-11	0	0	3	0	No barrier

CS <sub>2</sub> + hv → <sup>1</sup> CS + <sup>3</sup> S → <sup>1</sup> CS + <sup>1</sup> S		>223 nm 0% 0%	<223 nm 40% 60%		<a href="http://satellite.mpic.de/spectral_atlas">http://satellite.mpic.de/spectral_atlas</a> (Yang <i>et al.</i> 1980, Moltzen <i>et al.</i> 1988)	
OCS + hv → <sup>1</sup> CO + S( <sup>1</sup> D)		>291 nm 0%	<291 nm 100%		<a href="http://phidrates.space.swri.edu/">http://phidrates.space.swri.edu/</a>	
H <sub>2</sub> S + hv → HS + H → <sup>3</sup> S + H <sub>2</sub> → S( <sup>1</sup> D) + H <sub>2</sub>		>316 nm 0% 0% 0%	316-200 nm 100% 0% 0%	<200 nm 95% 3% 2%		<a href="http://phidrates.space.swri.edu/">http://phidrates.space.swri.edu/</a> (Feng <i>et al.</i> 1999, Watanabe & Jursa 1964, Hsu <i>et al.</i> 1992)
CS + hv → <sup>3</sup> S + <sup>3</sup> C		>168 nm 0%	<168 nm 100%			
HCS + hv → H + CS		>604 nm 0%	<604 nm 100%		EOM-CCSD. Some excited state may be not dissociative.	
H <sub>2</sub> CS + hv → H + HCS → H <sub>2</sub> + CS		>312 nm 0% 0%	<312 nm 100% 0%		(Judge <i>et al.</i> 1978) + EOM-CCSD	
CH <sub>3</sub> SH + hv → CH <sub>3</sub> + SH → CH <sub>3</sub> S + H	306 356	> 334nm 0% 0%	334-222 nm 100% 0%	222-190 nm 100-(222-λ)/0.64 (222-λ)/0.64	< 190 nm 50% 50%	EOM-CCSD + (Steer & Knight 1968, Vaghjiani 1993) Absorption: <a href="http://phidrates.space.swri.edu/">http://phidrates.space.swri.edu/</a> but branching ratio from <a href="http://phidrates.space.swri.edu/">http://phidrates.space.swri.edu/</a> above 250 nm likely wrong. The selective S-H bond fission at 248 nm in CH <sub>3</sub> SH photolysis is explained by assuming the Franck-Condon overlap between the ground vibrational wave function (at extended S-H geometries) and the dissociative surface to be on the CH <sub>3</sub> S+H exit channel side only (Vaghjiani 1993).
CH <sub>3</sub> CHS + hv → CH <sub>3</sub> + HCS → CH <sub>3</sub> CS + H	359 376	>318 nm 0% 0%	<318 nm 100% 0%		We neglect the CH <sub>3</sub> CS formation considering than H + CH <sub>3</sub> CS give CH <sub>3</sub> + HCS.	
		>274 nm	<274 nm		EOM-CCSD	

$C_2S + hv$	$\rightarrow C + CS$ $\rightarrow C_2 + S$	436 647	0% 0%	100% 0%			
$HCCS + hv$	$\rightarrow H + CCS$ $\rightarrow C_2H + S$ $\rightarrow CH + CS$	425 477 531	>281 nm 0% 0% 0%	282-251 nm 100% 0% 0%	<251 nm 80% 20% 0%	EOM-CCSD	
$H_2CCS + hv$	$\rightarrow H + HCCS$ $\rightarrow H_2 + HCCS$ $\rightarrow {}^3CH_2 + CS$ $\rightarrow {}^1CH_2 + CS$	375 372 475 513	>319 nm 0% 0% 0% 0%	<319 nm 100% 0% 0% 0%		EOM-CCSD	
$C_3S + hv$	$\rightarrow C_2 + CS$ $\rightarrow C_3 + {}^3S$ $\rightarrow C_3 + {}^1S$ $\rightarrow C + C_2S$	655 519 669 699	>182 nm 0% 0% 0% 0%	182-179nm 100% 0% 0% 0%	179-171nm 50% 0% 50% 0%	<171nm 40% 0% 40% 20%	EOM-CCSD
$H_2C_3S + hv$	$\rightarrow C_2H_2 + CS$ $\rightarrow H + HC_3S$ $\rightarrow H_2 + C_3S$	181 355 237	>290 nm 0% 0% 0%	<290 nm 30% 40% 30%		EOM-CCSD	
$H_2C_5S + hv$	$\rightarrow C_4H_2 + CS$ $\rightarrow H + HC_5S$ $\rightarrow C_2H_2 + C_3S$ $\rightarrow H_2 + C_5S$		>290 nm 0% 0% 0% 0%	<290 nm 50% 0% 50% 0%		Same absorption than $H_2C_3S$ . H + $HC_5S$ and $H_2 + C_5S$ very likely important channels but should give CS and $C_3S$ through photodissociation and reactions with H atoms.	
$NS + hv$	$\rightarrow N + S$	464	>258 nm 0%	<258 nm 100%			
$N_2S + hv$	$\rightarrow N_2 + S({}^1D)$	187	>640 nm 0%	<640 nm 100%		EOM-CCSD	
$HNCS + hv$	$\rightarrow H + NCS$ $\rightarrow HNC + S$	375 335	>320 nm 0% 0%	<320 nm 30% 70%		EOM-CCSD + (Boxall & Simons 1973)	

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